

Towards first order causal relativistic spin hydrodynamics

Samapan Bhadury.

Institute of Theoretical Physics, Jagiellonian University.

At

65th Cracow School of Theoretical Physics, Krakow, Poland.

June 17, 2025



Based On : [PRC 111 \(2025\), 034909](#) and [arXiv: 2503.08428](#)

Statement of Problems

Relativistic Kinetic Theory :

Relativistic Spin Hydrodynamics with ERTA :

Relativistic Spin hydrodynamics with NRTA :

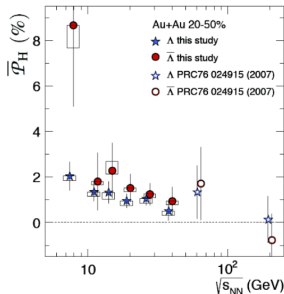
Summary and Outlook :

Particle Polarization :

- Spin polarization of hadrons in heavy-ion collisions, predicted in 2004.

[Z. T. Liang, X. N. Wang, PRL **94**, 102301 (2005), PLB **629**, 20 (2005)]

STAR Collaboration, *Global Lambda hyperon polarization in nuclear collisions*, Nature 548 62-65, 2017



First evidence of a quantum effect in (relativistic) hydrodynamics



Experimental evidence, [STAR Collaboration, Nature 548, 62 (2017), PRL 123, 132301 (2019), PRL 126, 162301 (2021)]

Theoretical models assuming equilibration of spin d.o.f. explains the data.

Particle Polarization :

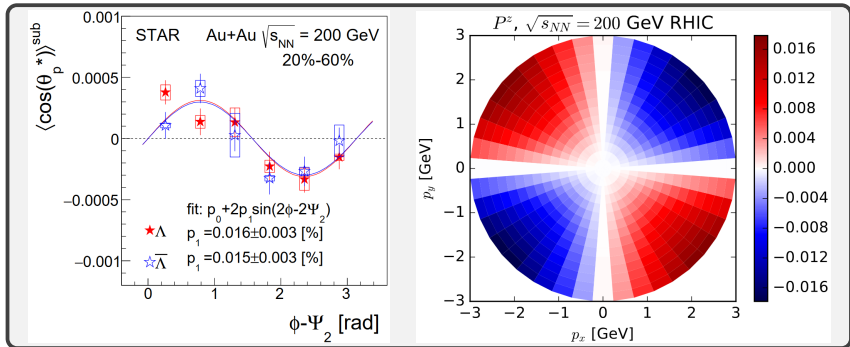


Figure 1: Observation (L) and prediction (R) of longitudinal polarization.

[Left: Phys. Rev. Lett. **123** 132301 (2019); Right: Phys. Rev. Lett. **120** 012302 (2018)]

- Inclusion of shear-induced polarization (SIP) solves the problem with extra constraints.
[Fu et. al. Phys. Rev. Lett. 127, 142301 (2021); Becattini et. al. Phys. Lett. B 820 136519 (2021)]
- Still the resolution remains ambiguous.
[Florkowski et. al., Phys. Rev. C 100, 054907 (2019); Phys.Rev.C 105, 064901 (2022)]
- Do dissipative forces play any role and solve the problem? [Sapna et. al., arXiv:2503.22552]

Spin NRTA Transport :

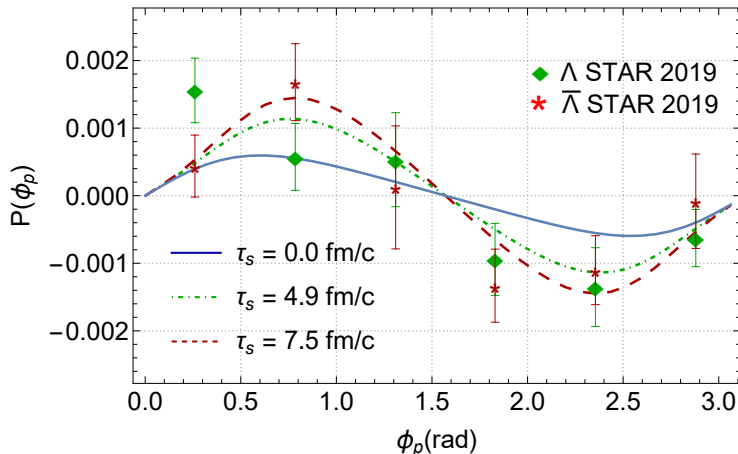


Figure 2: Polarization. $\tau_s = 7.5$ fm for $\bar{\Lambda}$ and, $\tau_s = 4.9$ fm for Λ . [S. Banerjee *et. al.*, accepted in PRC]

- τ_s is in agreement with [Hidaka *et. al.*, PRC **109** (2024), 054909, Wagner *et. al.*, PRR **6**, (2024) 043103].
- Incorporation of dissipative effects is necessary.

Relativistic Dissipative Spin Hydrodynamics :

- We first note that spin-polarization originates from the rotation of fluid.
- Hence, we will have to deal with three conserved currents :

$$\partial_\mu N^\mu = 0, \quad \partial_\mu T^{\mu\nu} = 0, \quad \partial_\lambda J^{\lambda,\mu\nu} = 0$$

where, $J = L + S$. Also, $L^{\lambda,\mu\nu} = x^\mu T^{\lambda\nu} - x^\nu T^{\lambda\mu}$.

[See talk by Zbigniew Drogosz]

- For symmetric $T^{\mu\nu}$ we have,

$$\partial_\lambda S^{\lambda,\mu\nu} = 0$$

$$N^\mu = N_{\text{eq}}^\mu + \delta N^\mu, \quad T^{\mu\nu} = T_{\text{eq}}^{\mu\nu} + \delta T^{\mu\nu}, \quad S^{\lambda,\mu\nu} = S_{\text{eq}}^{\lambda,\mu\nu} + \delta S^{\lambda,\mu\nu}$$

- The dissipative parts require microscopic description → **Kinetic Theory**.

Relativistic Dissipative Spin Hydrodynamics :

- Hydrodynamics is derived through an order-by-order gradient expansion:

$$X^{\mu_1 \cdots \mu_\ell} = X_{(0)}^{\mu_1 \cdots \mu_\ell} + X_{(1)}^{\mu_1 \cdots \mu_\ell} + \cdots + X_{(n)}^{\mu_1 \cdots \mu_\ell}$$

$$\begin{array}{ll} n = 0 \longrightarrow & \text{Perfect fluid dynamics.} \\ n = 1 \longrightarrow & \text{Navier - Stokes.} \\ n = 2 \longrightarrow & \text{MIS - type theories.} \end{array} \left. \vphantom{\begin{array}{l} n = 0 \\ n = 1 \\ n = 2 \end{array}} \right\} \text{Dissipative}$$

- Navier-Stokes theories lead to acausal theories.

[W. A. Hiscock et. al., *Annals Phys.* **151** (1983) 466-496, *PRD* **31** (1985) 725-733, **35** (1987) 3723-3732]

- MIS theories introduce extra degrees of freedom.

- Casual first-order theories require generalized frame definition.

[F. Bemfica et. al., *PRD* **98** (2018), 104064 *PRL* **122** (2019), 221602,

P. Kovtun *JHEP* **10** (2019) 034, *PRD* **106** (2022), 066023]

Summary of the Problem :

- Construct relativistic spin hydrodynamics in general frame.
 - With Extended RTA (ERTA).
 - With Novel RTA (NRTA).

Statement of Problems

Relativistic Kinetic Theory :

Relativistic Spin Hydrodynamics with ERTA :

Relativistic Spin hydrodynamics with NRTA :

Summary and Outlook :

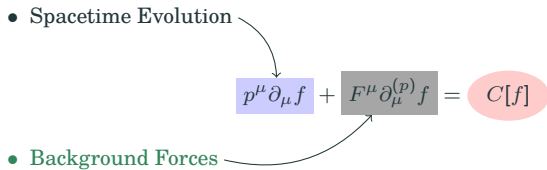
Boltzmann Equation

- Spacetime Evolution


$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

Boltzmann Equation

- Spacetime Evolution



The diagram shows the Boltzmann equation $p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$. An arrow from the 'Spacetime Evolution' bullet points to the first term $p^\mu \partial_\mu f$, which is enclosed in a light blue box. Another arrow from the 'Background Forces' bullet points to the second term $F^\mu \partial_\mu^{(p)} f$, which is enclosed in a gray box. The right-hand side $C[f]$ is enclosed in a light red oval.

$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

- Background Forces

Boltzmann Equation

- Spacetime Evolution

The diagram shows the Boltzmann equation: $p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$. Arrows indicate the physical meaning of each term: an arrow from 'Spacetime Evolution' points to the first term; an arrow from 'Background Forces' points to the second term; and an arrow from 'Collision Kernel' points to the right-hand side.

$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

- Background Forces
- Collision Kernel

Boltzmann Equation

- Spacetime Evolution

The diagram shows the Boltzmann equation: $p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$. Arrows indicate the physical meaning of each term: an arrow from 'Spacetime Evolution' points to the first term; an arrow from 'Background Forces' points to the second term; and an arrow from 'Collision Kernel' points to the right-hand side.

$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

- Background Forces
- Collision Kernel

Boltzmann Equation

- Spacetime Evolution

The diagram shows the Boltzmann equation: $p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$. An arrow points from 'Spacetime Evolution' to the first term $p^\mu \partial_\mu f$. Another arrow points from 'Background Forces' to the second term $F^\mu \partial_\mu^{(p)} f$. A third arrow points from 'Collision Kernel' to the right-hand side $C[f]$.

$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

- Background Forces
- Collision Kernel

- Gravitational Forces : $F^\mu = -\Gamma_{\alpha\beta}^\mu p^\alpha p^\beta.$
- Electromagnetic Forces : $F^\mu = q F^{\mu\nu} p_\nu.$
- Mean-Field Forces : $F^\mu = M (\partial^\mu M).$

Boltzmann Equation

- Spacetime Evolution

$$p^\mu \partial_\mu f + F^\mu \partial_\mu^{(p)} f = C[f]$$

The diagram shows the Boltzmann equation with three colored boxes: a blue box for $p^\mu \partial_\mu f$, a grey box for $F^\mu \partial_\mu^{(p)} f$, and a red oval for $C[f]$. Arrows point from the text labels to these boxes: 'Spacetime Evolution' to the blue box, 'Background Forces' to the grey box, and 'Collision Kernel' to the red oval.

- Background Forces
- Collision Kernel

- Gravitational Forces : $F^\mu = -\Gamma_{\alpha\beta}^\mu p^\alpha p^\beta$.
- Electromagnetic Forces : $F^\mu = q F^{\mu\nu} p_\nu$.
- Mean-Field Forces : $F^\mu = M (\partial^\mu M)$.

- Global Equilibrium solution:

$$f_0 = (\exp g + r)^{-1}$$

where, $r = 0, \pm 1$ and $g \equiv \sum_n \alpha_n \phi_n$

- Under local equilibrium, $\alpha_n \rightarrow \alpha_n(x^\mu)$

- For $2 \leftrightarrow 2$ collisions:

$$C[f] = \int dP dP' dK' \underbrace{\mathcal{W}_{\mathbf{k}\mathbf{k}' \leftrightarrow \mathbf{p}\mathbf{p}'}}_{\text{Transition Amplitude}} \times (f_{\mathbf{p}} f_{\mathbf{p}'} - f_{\mathbf{k}} f_{\mathbf{k}'})$$

Transition Amplitude:



- For $2 \leftrightarrow 2$ collisions:

$$C[f] = \int dP dP' dK' \underbrace{\mathcal{W}_{\mathbf{k}\mathbf{k}' \leftrightarrow \mathbf{p}\mathbf{p}'}}_{\text{Transition Amplitude}} \times (f_{\mathbf{p}}f_{\mathbf{p}'} - f_{\mathbf{k}}f_{\mathbf{k}'})$$

Transition Amplitude:

- Chapman-Enskog Expansion:

$$f_{\mathbf{k}} = f_{0\mathbf{k}} + \delta f_{\mathbf{k}} = f_{0\mathbf{k}} (1 + \phi_{\mathbf{k}})$$

The Collision Kernel

- For $2 \leftrightarrow 2$ collisions:

$$C[f] = \int dP dP' dK' \underbrace{W_{\mathbf{k}\mathbf{k}' \leftrightarrow \mathbf{p}\mathbf{p}'}}_{\text{Transition Amplitude}} \times (f_{\mathbf{p}} f_{\mathbf{p}'} - f_{\mathbf{k}} f_{\mathbf{k}'})$$

Transition Amplitude:

- Chapman-Enskog Expansion:

$$f_{\mathbf{k}} = f_{0\mathbf{k}} + \delta f_{\mathbf{k}} = f_{0\mathbf{k}} (1 + \phi_{\mathbf{k}})$$

- Linearized Collision Operator :

$$C[f] \rightarrow \hat{L}\phi_{\mathbf{k}} = \int dP dP' dK' W_{\mathbf{k}\mathbf{k}' \leftrightarrow \mathbf{p}\mathbf{p}'} f_{0\mathbf{k}} f_{0\mathbf{k}'} \times (\phi_{\mathbf{p}} + \phi_{\mathbf{p}'} - \phi_{\mathbf{k}} - \phi_{\mathbf{k}'})$$

Collisional Invariants:

- Collisional invariants remain conserved during collisions.
- Each collisional invariant correspond to a conservation law.
- For a non-rotating, unpolarizable fluid :
 - $\phi \equiv 1 \longrightarrow$ Number Conservation.
 - $\phi \equiv E_{\mathbf{k}} \longrightarrow$ Energy Conservation.
 - $\phi \equiv \vec{k} (\sim k^{\langle \mu \rangle}) \longrightarrow$ Linear Momentum Conservation.

- Thus, a collision kernel should satisfy:

$$\hat{L} 1 = 0, \quad \hat{L} E_{\mathbf{k}} = 0, \quad \hat{L} k^{\langle \mu \rangle} = 0.$$

- The linearized collision kernel satisfies the property:

$$\begin{aligned} \int dK \psi_{\mathbf{k}} \hat{L} \phi_{\mathbf{k}} &= \int dK \phi_{\mathbf{k}} \hat{L} \psi_{\mathbf{k}} \\ \implies \int dK \hat{L} \phi_{\mathbf{k}} &= 0, \quad \int dK k^{\mu} \hat{L} \phi_{\mathbf{k}} = \int dK \left(u^{\mu} E_{\mathbf{k}} + k^{\langle \mu \rangle} \right) \hat{L} \phi_{\mathbf{k}} = 0. \end{aligned}$$

- We will work with two types linearized collision kernels.
 1. Extended Relaxation Time Approximation (ERTA):

$$\hat{L}_{\text{ERTA}} \phi_{\mathbf{k}} = -\frac{E_{\mathbf{k}}}{\tau_{\text{R}}} (\phi_{\mathbf{k}} - \phi_{\mathbf{k}}^*) f_{0\mathbf{k}}$$

where,

$$\phi_{1,\mathbf{k}}^* = -\frac{(k \cdot \delta u)}{T} + \frac{(E_{\mathbf{k}} - \mu) \delta T}{T^2} + \frac{\delta \mu}{T}$$
$$\delta u_{\mu} = u_{\mu}^* - u_{\mu}, \quad \delta T = T^* - T, \quad \delta \mu = \mu^* - \mu.$$

[D. Dash *et. al.*, PLB 831 (2022) 137202]

- We will work with two types linearized collision kernels.
 1. **Extended Relaxation Time Approximation (ERTA):**

$$\hat{L}_{\text{ERTA}} \phi_{\mathbf{k}} = -\frac{E_{\mathbf{k}}}{\tau_{\text{R}}} (\phi_{\mathbf{k}} - \phi_{\mathbf{k}}^*) f_{0\mathbf{k}}$$

where,

$$\phi_{1,\mathbf{k}}^* = -\frac{(k \cdot \delta u)}{T} + \frac{(E_{\mathbf{k}} - \mu) \delta T}{T^2} + \frac{\delta \mu}{T}$$
$$\delta u_{\mu} = u_{\mu}^* - u_{\mu}, \quad \delta T = T^* - T, \quad \delta \mu = \mu^* - \mu.$$

[D. Dash *et. al.*, PLB 831 (2022) 137202]

2. **Novel Relaxation Time Approximation (NRTA):**

$$\hat{L}_{\text{NRTA}} \sim \left(-\mathbb{1} + \sum_{n=1}^5 |\lambda_n\rangle \langle \lambda_n| \right)$$

where, $|\lambda_n\rangle$ are degenerate, orthogonal eigenvectors of $\hat{L}_{\text{NRTA}}$.

[G. S. Rocha *et. al.*, PRL 127 (2021), 042301]

Solving The Boltzmann Equation

- Boltzmann equation is an inhomogeneous partial differential equation.

Solving The Boltzmann Equation

- Boltzmann equation is an inhomogeneous partial differential equation.
- Thus, the solution is:

$$\phi = \phi_h + \phi_{ih}$$

- The homogeneous part looks like :

$$\phi_h = a + b_\mu k^\mu = a \cdot 1 + (b \cdot u) E_{\mathbf{k}} + b_{\langle \mu \rangle} k^{\langle \mu \rangle}.$$

- The coefficients, a, b_μ can be determined from frame and matching conditions.
- Subtracting the homogeneous part of the solution, gives freedom to choose the frame and matching conditions.

[S. R. de Groot *et. al.*, Relativistic Kinetic Theory, C. Cercignani *et. al.*, The Relativistic Boltzmann Equation]

Solving The Boltzmann Equation

- Boltzmann equation is an inhomogeneous partial differential equation.
- Thus, the solution is:

$$\phi = \phi_h + \phi_{ih}$$

- The homogeneous part looks like :

$$\phi_h = a + b_\mu k^\mu = a \cdot 1 + (b \cdot u) E_{\mathbf{k}} + b_{\langle \mu \rangle} k^{\langle \mu \rangle}.$$

- The coefficients, a, b_μ can be determined from frame and matching conditions.
- Subtracting the homogeneous part of the solution, gives freedom to choose the frame and matching conditions.

[S. R. de Groot *et. al.*, Relativistic Kinetic Theory, C. Cercignani *et. al.*, The Relativistic Boltzmann Equation]

- Both ERTA and NRTA have this feature.

$$\text{ERTA :} \quad a \rightarrow (E_{\mathbf{k}} - \mu) (\delta T / T^2) + (\delta \mu / T), \quad \text{and,} \quad b_\mu \rightarrow -(\delta u_\mu / T).$$

$$\text{NRTA :} \quad a \rightarrow \Phi_0, \quad \text{and,} \quad b_\mu \rightarrow \Phi_0^{\langle \mu \rangle}.$$

[D. Dash *et. al.*, PLB 831 (2022) 137202, G. S. Rocha *et. al.*, PRL 127 (2021), 042301]

Solving The Boltzmann Equation (Contd.)

- Two popular approaches are considered :

- Chapman-Enskog-like iterative solution:

$$\phi_{n,\mathbf{k}} f_{0\mathbf{k}} = \phi_{n,\mathbf{k}}^* f_{0\mathbf{k}} - \left(\frac{\tau_R}{E_{\mathbf{k}}} \right) (k \cdot \partial) f_{(n-1)\mathbf{k}},$$

We will use this to solve the Extended RTA case.

[D. Dash *et. al.*, PLB 831 (2022) 137202]

- Moment method:

$$\phi_{\mathbf{k}} = \sum_{n,\ell=0}^{\infty} \Phi_n^{\langle \mu_1 \dots \mu_\ell \rangle} k_{\langle \mu_1} \dots k_{\mu_\ell \rangle} P_n^{(\ell)}(\beta E_{\mathbf{k}})$$

We will use this to solve the Novel RTA case. Here $P_n^{(\ell)}$ are orthogonal polynomials satisfying the property:

$$\frac{\ell!}{(2\ell+1)!!} \left\langle (E_{\mathbf{k}}/\tau_R) (k \cdot \Delta \cdot k)^\ell P_n^{(\ell)} P_m^{(\ell)} \right\rangle_{0\mathbf{k}} = A_n^{(\ell)} \delta_{nm},$$
$$A_n^{(\ell)} = \frac{\ell!}{(2\ell+1)!!} \left\langle (E_{\mathbf{k}}/\tau_R) (k \cdot \Delta \cdot k)^\ell P_n^{(\ell)} P_n^{(\ell)} \right\rangle_{0\mathbf{k}}$$

[G. S. Rocha *et. al.*, PRL 127 (2021), 042301]

- Phase-space is extended to include spin degrees of freedom :

$$f_{\mathbf{k}}(x, k) \longrightarrow f_s(x, k, s)$$

$$f_{0\mathbf{k}} \longrightarrow f_{0,s} = f_{0\mathbf{k}} \exp(s : \omega) \approx f_{0\mathbf{k}} \left[1 + \frac{1}{2} (s : \omega) \right] + \mathcal{O}(\omega^2)$$

- Homogeneous part for spin-polarizable particles :

$$\phi_h = a + b_\mu k^\mu + c_{\mu\nu} s^{\mu\nu}$$

- The solutions are modified as :

- Chapman-Enskog-like iterative solution (ERTA):

$$\phi_{n,s} f_{0,s} = \phi_{n,s}^* f_{0,s} - \frac{\tau_R}{(u \cdot p)} (p \cdot \partial) f_{(n-1),s},$$

- Moment method (NRTA):

$$\phi_s = \sum_{n,\ell=0}^{\infty} \left(\Phi_n^{\langle \mu_1 \cdots \mu_\ell \rangle} + s_{\mu\nu} \Psi_n^{\mu\nu, \langle \mu_1 \cdots \mu_\ell \rangle} \right) k_{\langle \mu_1} \cdots k_{\mu_\ell \rangle} P_n^{(\ell)}(\beta E_{\mathbf{k}})$$

Section Outline :

Statement of Problems

Relativistic Kinetic Theory :

Relativistic Spin Hydrodynamics with ERTA :

Relativistic Spin hydrodynamics with NRTA :

Summary and Outlook :

- The thermodynamic variables (starred) are determined via field re-definitions :

$$\begin{aligned} \langle q_1 \phi_s \rangle_o + \langle \bar{q}_1 \bar{\phi}_s \rangle_{\bar{o}} &= 0, & \langle q_2 \phi_s \rangle_o + \langle \bar{q}_2 \bar{\phi}_s \rangle_{\bar{o}} &= 0 \\ \langle q_3 k^{\langle \mu \rangle} \phi_s \rangle_o + \langle \bar{q}_3 k^{\langle \mu \rangle} \bar{\phi}_s \rangle_{\bar{o}} &= 0, & \langle q_4 s^{\mu\nu} \phi_s \rangle_o + \langle \bar{q}_4 s^{\mu\nu} \bar{\phi}_s \rangle_{\bar{o}} &= 0 \end{aligned}$$

where we use the notations:

$$\langle (\cdots) \rangle_o = \int dK dS (\cdots) f_{o\mathbf{k}}, \quad \langle (\cdots) \rangle_{\bar{o}} = \int dK dS (\cdots) \bar{f}_{o\mathbf{k}}$$

- The thermodynamic variables are :

$$X^* = X + \delta X$$

where, $X \equiv T, \mu, u_\mu, \omega_{\mu\nu}$.

- Then we find using the field-redefinitions (up to $\mathcal{O}(\partial)$):

$$\begin{aligned} \delta u_\mu &= \beta \mathcal{C}_1 (\nabla_\mu \xi), & \delta T &= \mathcal{C}_2 \theta, & \delta \mu &= \mathcal{C}_3 \theta, \\ \delta \omega^{\mu\nu} &= \mathcal{D}_H^{\mu\nu} \theta + \mathcal{D}_n^{\mu\nu\gamma} (\nabla_\gamma \xi) + \mathcal{D}_\pi^{\mu\nu\alpha\beta} \sigma_{\alpha\beta} + \mathcal{D}_\Sigma^{\mu\nu\gamma\alpha\beta} (\nabla_\gamma \omega_{\alpha\beta}). \end{aligned}$$

- Even with arbitrary frame and matching conditions, we had to use, $D \rightarrow \nabla$.
- While the first-order theory is still acausal, we can now have, $\tau_R(x, p, s)$.

- Under a general field re-definition the constitutive relations are given by :

$$\begin{aligned} N^\mu &= (n_o + \delta n) u^\mu + n^\mu, \\ T^{\mu\nu} &= (\mathcal{E}_o + \delta\mathcal{E}) u^\mu u^\nu - (\mathcal{P} + \delta\mathcal{P}) \Delta^{\mu\nu} + 2h^{(\mu} u^{\nu)} + \pi^{\mu\nu}, \\ S^{\lambda,\mu\nu} &= S_o^{\lambda,\mu\nu} + \delta S^{\lambda,\mu\nu}. \end{aligned}$$

- The dissipative currents are :

$$\begin{aligned} \delta n &= \nu \theta & \delta\mathcal{E} &= e \theta, & \delta\mathcal{P} &= \rho \theta, \\ n^\mu &= \kappa_n^{\mu\nu} (\nabla_\nu \xi), & h^\mu &= \kappa_h^{\mu\nu} (\nabla_\nu \xi), \\ \pi^{\mu\nu} &= 2\eta \sigma^{\mu\nu}, \\ \delta S^{\lambda,\mu\nu} &= B_{II}^{\mu\nu} \theta + B_n^{\mu\nu\gamma} (\nabla_\gamma \xi) + B_\pi^{\mu\nu\alpha\beta} \sigma_{\alpha\beta} + B_\Sigma^{\mu\nu\gamma\alpha\beta} (\nabla_\gamma \omega_{\alpha\beta}) \end{aligned}$$

- The expressions of the transport coefficients can be obtained assuming:

$$\tau_R(x, p, s) = \tau_{\text{eq}}(x) (\beta \cdot p)^{\ell_1} (u \cdot s)^{\ell_2}$$

- The entropy production is given by,

$$\partial_\mu \mathcal{H}_\mu = -\beta \Pi \theta - \mathcal{Q}^\mu (\nabla_\mu \xi) + \beta \pi^{\mu\nu} \sigma_{\mu\nu} - \mathcal{S}^{\lambda\mu\nu} (\nabla_\lambda \omega_{\mu\nu})$$

- The frame-invariant transport coefficients are:

$$\Pi = -\zeta \theta = \delta \mathcal{P} - \left(\frac{\chi_b}{\beta} \right) \delta \mathcal{E} + \left(\frac{\chi_a}{\beta} \right) \delta n,$$

$$\mathcal{Q}^\mu = \kappa^{\mu\nu} (\nabla_\nu \xi) = n^\mu - \left(\frac{n_o}{\mathcal{E}_o + \mathcal{P}_o} \right) h^\mu,$$

$$\pi^{\mu\nu} = 2\eta \sigma^{\mu\nu},$$

$$\mathcal{S}^{\lambda\mu\nu} = \beta \Sigma^{\lambda\mu\nu\gamma\alpha\beta} (\nabla_\gamma \omega_{\alpha\beta}) = \frac{1}{2} \left(u^\rho D_\Sigma^{\alpha\beta\lambda\mu\nu} \delta S_{\rho,\alpha\beta} - \delta S^{\lambda,\mu\nu} \right)$$

[S. B., PRC 111 (2025), 034909]

Section Outline :

Statement of Problems

Relativistic Kinetic Theory :

Relativistic Spin Hydrodynamics with ERTA :

Relativistic Spin hydrodynamics with NRTA :

Summary and Outlook :

The Linearized collision kernel - NRTA

- We work with $\mu = 0$, as NRTA is not built for pair production and annihilation.

- We have 10 collisional invariants and hence : $\hat{L}\phi_s \equiv -1 + \sum_{n=1}^{10} |\lambda_n\rangle \langle \lambda_n | \phi_s \rangle$

- The orthonormal basis are :

$$|\lambda_1\rangle = \frac{E_{\mathbf{k}}}{\sqrt{\langle (E_{\mathbf{k}}^3/\tau_R) \rangle_o}},$$

$$|\lambda_{2-4}\rangle = \frac{k^{\langle\mu\rangle}}{\sqrt{(1/3) \langle (E_{\mathbf{k}}/\tau_R) k_{\langle\alpha\rangle} k^{\langle\alpha\rangle} \rangle_o}},$$

$$|\lambda_{5-7}\rangle = \sqrt{\frac{3}{\langle (E_{\mathbf{k}}/\tau_R)(\tilde{s} \cdot \tilde{s}) \rangle_o}} \left[\tilde{s}^\mu - \frac{\langle (E_{\mathbf{k}}^2/\tau_R) \tilde{s}^\mu \rangle_o}{\langle (E_{\mathbf{k}}^3/\tau_R) \rangle_o} E_{\mathbf{k}} - \frac{\langle (E_{\mathbf{k}}/\tau_R) k^{\langle\gamma\rangle} \tilde{s}^\mu \rangle_o k_{\langle\gamma\rangle}}{\langle (1/3) (E_{\mathbf{k}}/\tau_R) k^{\langle\alpha\rangle} k_{\langle\alpha\rangle} \rangle_o} \right],$$

$$|\lambda_{8-10}\rangle = \sqrt{\frac{3}{\langle (E_{\mathbf{k}}/\tau_R)(\tilde{s} : \tilde{s}) \rangle_o}} \left[\tilde{s}^{\mu\nu} - \frac{\langle (E_{\mathbf{k}}^2/\tau_R) \tilde{s}^{\mu\nu} \rangle_o}{\langle (E_{\mathbf{k}}^3/\tau_R) \rangle_o} E_{\mathbf{k}} - \frac{\langle (E_{\mathbf{k}}/\tau_R) k^{\langle\mu\rangle} \tilde{s}^{\mu\nu} \rangle_o k_{\langle\mu\rangle}}{\langle (1/3) (E_{\mathbf{k}}/\tau_R) k^{\langle\alpha\rangle} k_{\langle\alpha\rangle} \rangle_o} \right].$$

where, $\tilde{s}^\mu = \Delta^{\mu\alpha} u^\beta s_{\alpha\beta}$, and $\tilde{s}^{\mu\nu} = \Delta^{\mu\alpha} \Delta^{\nu\beta} s_{\alpha\beta}$.

- Then we get,

$$\hat{L}E_{\mathbf{k}} = 0, \quad \hat{L}k^{\langle\mu\rangle} = 0, \quad \hat{L}s^{\mu\nu} = 0.$$

Field Redefinition - NRTA

- We work with $\mu = 0$, as NRTA is not built for pair production and annihilation.
- Now we use a new set of notations :

$$\langle(\cdots)\rangle_o = \int dK dS (\cdots) f_{oS}, \quad \langle(\cdots)\rangle_{o\mathbf{k}} = \int dK (\cdots) f_{o\mathbf{k}}$$

- The thermodynamic variables (starred) are determined via field redefinitions:

$$\int dK dS q_1 \phi_s f_{o\mathbf{k}} = 0, \quad \int dK dS q_2 k^{\langle\mu\rangle} \phi_s f_{o\mathbf{k}} = 0, \quad \int dK dS q_3 s^{\mu\nu} \phi_s f_{o\mathbf{k}} = 0$$

- The solution takes the form:

$$\phi_s = \phi_{\mathbf{p}} + s : \psi_{\mathbf{p}}.$$

with

$$\phi_{\mathbf{p}} = \sum_{n \in S_0^{(\ell)}} \sum_{\ell=0}^{\infty} \Phi_n^{\langle\mu_1 \cdots \mu_{\ell}\rangle} p_{\langle\mu_1} \cdots p_{\mu_{\ell}\rangle} P_n^{(0,\ell)}, \quad \text{and,} \quad \psi_{\mathbf{p}}^{\mu\nu} = \sum_{n \in S_1^{(\ell)}} \sum_{\ell=0}^{\infty} \Psi_n^{\mu\nu, \langle\mu_1 \cdots \mu_{\ell}\rangle} p_{\langle\mu_1} \cdots p_{\mu_{\ell}\rangle} P_n^{(1,\ell)},$$

$$\frac{\ell!}{(2\ell+1)!!} \int dP (E_{\mathbf{p}}/\tau_R) (p \cdot \Delta \cdot p)^{\ell} P_m^{(j,\ell)} P_n^{(j,\ell)} f_{o\mathbf{p}} = A_n^{(j,\ell)} \delta_{mn}.$$

- The homogeneous parts of the solution are :

$$\begin{aligned}\Phi_0 &= - \sum_{n \in \mathbb{S}_0^{(0)} - \{0\}} \frac{\left\langle q_1 P_n^{(0,0)} \right\rangle_{\text{op}}}{\left\langle q_1 P_0^{(0,0)} \right\rangle_{\text{op}}} \Phi_n, & \Phi_0^{\langle \mu \rangle} &= - \frac{\left\langle q_2 (p \cdot \Delta \cdot p) P_1^{(0,1)} \right\rangle_{\text{op}}}{\left\langle q_2 (p \cdot \Delta \cdot p) P_0^{(0,1)} \right\rangle_{\text{op}}} \Phi_1^{\langle \mu \rangle}, \\ \Psi_0^{\mu\nu} &= - \left[\frac{\left\langle q_3 P_1^{(1,0)} \right\rangle_{\text{op}}}{\left\langle q_3 P_0^{(1,0)} \right\rangle_{\text{op}}} \right] \Psi_1^{\mu\nu}.\end{aligned}$$

- Using these, one can obtain the expressions of all $\Phi_n^{\langle \mu_1 \cdots \mu_\ell \rangle}$ and $\Psi_n^{\mu\nu, \langle \mu_1 \cdots \mu_\ell \rangle}$.

[S. B., 2503.08428]

Statement of Problems

Relativistic Kinetic Theory :

Relativistic Spin Hydrodynamics with ERTA :

Relativistic Spin hydrodynamics with NRTA :

Summary and Outlook :

Summary and Outlook :

- **Summary :**

1. ERTA provides a simple solution to make the relaxation time dependent on p and s .
2. ERTA allows us to have a general choice of frame and matching conditions.
3. ERTA does not lead to first-order causal theory of spin hydrodynamics.
4. NRTA gives the option of constructing first-order causal spin-hydrodynamics.
5. NRTA cannot describe a system with pair production and annihilation.

- **Outlook :**

1. Construction of spin-BDNK theory is required.
2. Perform a linear mode analysis to verify the causality and stability of the theory.

Thank you.

- The frame-invariant transport coefficients are:

$$f_i \equiv \pi_i - \left(\frac{\partial \mathcal{P}_o}{\partial \mathcal{E}_o} \right)_{n_o} \varepsilon_i + \left(\frac{\partial \mathcal{P}_o}{\partial n_o} \right)_{\varepsilon_o} \nu_i,$$
$$l_i \equiv \gamma_i - \left(\frac{n_o}{\mathcal{E}_o + \mathcal{P}_o} \right) \theta_i.$$

- Out of the 16 parameters, only three one-derivative transport coefficients $\zeta(f_1, f_2, f_3)$, η , $\kappa(l_1, l_2)$ are independent.

- Full collision kernel for Novel RTA with spin :

[S. B. [arXiv:2503.08428](https://arxiv.org/abs/2503.08428)]

$$\begin{aligned}
 \hat{L}\phi_s = & - \left(\frac{E_{\mathbf{p}}}{\tau_R} \right) f_{0s} \left\{ \phi_s - \frac{\langle (E_{\mathbf{p}}^2/\tau_R) \phi_s \rangle_0}{\langle (E_{\mathbf{p}}^3/\tau_R) \rangle_0} E_{\mathbf{p}} - \frac{\langle (E_{\mathbf{p}}/\tau_R) p^{(\mu)} \phi_s \rangle_0}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) p^{(\alpha)} p_{(\alpha)} \rangle_0} p_{(\mu)} \right. \\
 & - \left[\langle (E_{\mathbf{p}}/\tau_R) \tilde{s}_\mu \phi_s \rangle_0 - \frac{\langle (E_{\mathbf{p}}^2/\tau_R) \tilde{s}_\mu \rangle_0}{\langle (E_{\mathbf{p}}^3/\tau_R) \rangle_0} \langle (E_{\mathbf{p}}^2/\tau_R) \phi_s \rangle_0 - \frac{\langle (E_{\mathbf{p}}/\tau_R) p^{(\gamma)} \tilde{s}_\mu \rangle_0}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) p^{(\alpha)} p_{(\alpha)} \rangle_0} \langle (E_{\mathbf{p}}/\tau_R) p_{(\gamma)} \phi_s \rangle_0 \right] \\
 & \times \left[\tilde{s}^\mu - \frac{\langle (E_{\mathbf{p}}^2/\tau_R) \tilde{s}^\mu \rangle_0}{\langle (E_{\mathbf{p}}^3/\tau_R) \rangle_0} E_{\mathbf{p}} - \frac{\langle (E_{\mathbf{p}}/\tau_R) p^{(\rho)} \tilde{s}^\mu \rangle_0}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) p^{(\beta)} p_{(\beta)} \rangle_0} p_{(\rho)} \right] \frac{1}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) (\tilde{s} \cdot \tilde{s}) \rangle_0} \\
 & - \left[\langle (E_{\mathbf{p}}/\tau_R) \tilde{s}_{\mu\nu} \phi_s \rangle_0 - \frac{\langle (E_{\mathbf{p}}^2/\tau_R) \tilde{s}_{\mu\nu} \rangle_0}{\langle (E_{\mathbf{p}}^3/\tau_R) \rangle_0} \langle (E_{\mathbf{p}}^2/\tau_R) \phi_s \rangle_0 - \frac{\langle (E_{\mathbf{p}}/\tau_R) p^{(\gamma)} \tilde{s}_{\mu\nu} \rangle_0}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) p^{(\alpha)} p_{(\alpha)} \rangle_0} \langle (E_{\mathbf{p}}/\tau_R) p_{(\gamma)} \phi_s \rangle_0 \right] \\
 & \times \left[\tilde{s}^{\mu\nu} - \frac{\langle (E_{\mathbf{p}}^2/\tau_R) \tilde{s}^{\mu\nu} \rangle_0}{\langle (E_{\mathbf{p}}^3/\tau_R) \rangle_0} E_{\mathbf{p}} - \frac{\langle (E_{\mathbf{p}}/\tau_R) p^{(\rho)} \tilde{s}^{\mu\nu} \rangle_0}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) p^{(\beta)} p_{(\beta)} \rangle_0} p_{(\rho)} \right] \frac{1}{\langle (1/3) (E_{\mathbf{p}}/\tau_R) (\tilde{s} : \tilde{s}) \rangle_0} \Bigg\},
 \end{aligned}$$

Inclusion of Dissipation :

- Decomposition of conserved current :

$$T^{\mu\nu} = \varepsilon u^\mu u^\nu - P \Delta^{\mu\nu} + q^\mu u^\nu + q^\nu u^\mu + \pi^{\mu\nu} = \langle p^\mu p^\nu \rangle ,$$

$$S^{\lambda,\mu\nu} = u^\lambda S^{\mu\nu} + (u^\mu \Delta^{\nu\lambda} - u^\nu \Delta^{\mu\lambda}) \Sigma + u^\mu \Sigma_{(s)}^{(\lambda\nu)} - u^\nu \Sigma_{(s)}^{(\lambda\mu)} + u^\mu \Sigma_{(a)}^{\lambda\nu} - u^\nu \Sigma_{(a)}^{\lambda\mu} + \Sigma^{\lambda,\mu\nu} = \langle p^\lambda s^{\mu\nu} \rangle ,$$

- Transport Coefficients are given by :

$$\begin{aligned} \delta\varepsilon &= \sum_{n \in \mathbb{S}_0^{(0)}} \Phi_n \left\langle E_{\mathbf{p}}^2 P_n^{(0,0)} \right\rangle_{0\mathbf{p}} , & \delta P &= - \sum_{n \in \mathbb{S}_0^{(0)}} \Phi_n \left\langle (1/3) (p \cdot \Delta \cdot p) P_n^{(0,0)} \right\rangle_{0\mathbf{p}} , \\ q^\mu &= \sum_{n \in \mathbb{S}_0^{(1)}} \Phi_n^{(\mu_1)} \left\langle (1/3) E_{\mathbf{p}} (p \cdot \Delta \cdot p) P_n^{(0,1)} \right\rangle_{0\mathbf{p}} , & \pi^{\mu\nu} &= \Phi_0^{(\mu\nu)} \left\langle (2/15) (p \cdot \Delta \cdot p)^2 \right\rangle_{0\mathbf{p}} , \\ \delta S^{\mu\nu} &= \sum_{n \in \mathbb{S}_1^{(0)}} \Psi_n^{\mu\nu} \left\langle E_{\mathbf{p}} P_n^{(1,0)} \right\rangle_{0\mathbf{p}} , & \delta \Sigma_{(s)}^{(\mu\nu)} &= u_\alpha \Delta_{\rho\gamma}^{\mu\nu} \Psi_0^{\alpha\rho, \langle\gamma\rangle} \left\langle (1/3) (p \cdot \Delta \cdot p) \right\rangle_{0\mathbf{p}} , \\ \delta \Sigma_{(a)}^{\mu\nu} &= \sum_{n \in \mathbb{S}_1^{(0)}} u_\alpha \Psi_n^{\alpha[\mu} u^{\nu]} \left\langle E_{\mathbf{p}} P_n^{(1,0)} \right\rangle_{0\mathbf{p}} + u_\alpha \Psi_0^{\alpha[\mu, \langle\nu\rangle]} \left\langle (1/3) (p \cdot \Delta \cdot p) \right\rangle_{0\mathbf{p}} , & \delta \Sigma^{\lambda,\mu\nu} &= \Psi_0^{\langle\mu\rangle\langle\nu\rangle, \langle\lambda\rangle} \left\langle (p \cdot \Delta \cdot p) \right\rangle_{0\mathbf{p}} , \end{aligned}$$