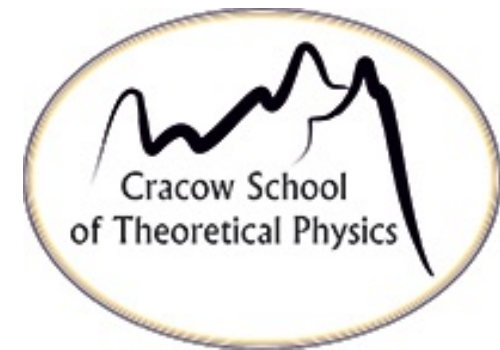


Loop Quantum Gravity and Quantum Information

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- Lec. 1 LQG, ref: Ashtekar-Bianchi, "A short review..." Rep. Prog. Phys 2021 [2104.04394]
- Lec. 2 LQG & QI, ref: Bianchi-Livine, in "Handbook of QG" Springer 2023 [2302.05922]




Zakopane
June 17, 2024



QISS

THE QUANTUM INFORMATION
STRUCTURE OF SPACETIME

■ Semiclassical regime of LQG

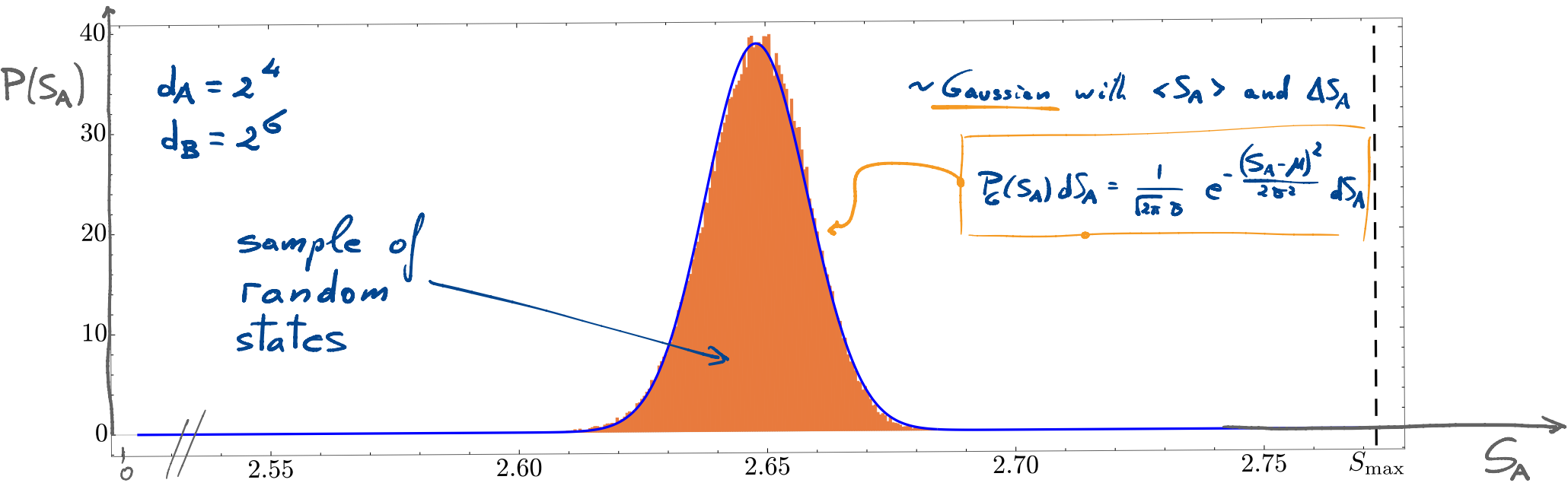
- Ground state? c.f. Hydrogen atom: $|1s\rangle$ vs. 
- Semiclassical states with fixed $\langle\psi|\hat{O}|\psi\rangle$, $\langle\psi|\hat{Q}_1\hat{Q}_2|\psi\rangle$, ...

■ Quantum information — new tools:

- semiclassical corner of the Hilbert space
- entanglement bounds on correlations
- subsystems from subalgebras
- causal origin of correlations

1 Random states and typical entanglement entropy

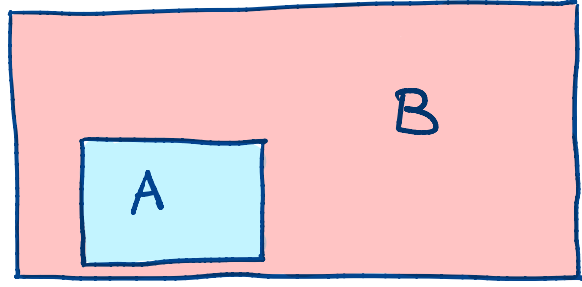
* Given a random pure state and a subsystem A, what is the probability of finding entanglement entropy S_A ?



Setup [Page PRL'1993]

• Hilbert space $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$, $\dim = d_A \times d_B$

• Random pure state $|\psi\rangle = U |\psi_0\rangle$
 uniform distribution: random unitary (Haar) \nearrow reference state



• Entanglement entropy $S_A(|\psi\rangle) = -\text{Tr}_A(\rho_A \log \rho_A)$, $\rho_A = \text{Tr}_B(|\psi\rangle\langle\psi|)$

□ Exact formulas for the average and the variance of $P(S_A)$

■ Average Entropy [Page, PRL'93] $1 < d_A \leq d_B$

$$\langle S_A \rangle = \Psi(d_A d_B + 1) - \Psi(d_B + 1) - \frac{d_A - 1}{2 d_B}$$

Gamma funct. $\Gamma(x)$
 diGamma funct. $\Psi(x) = \frac{\Gamma'(x)}{\Gamma(x)}$

• Asymptotics: $\langle S_A \rangle \sim \log d_A$ average entropy \approx max entropy
 $d_B \gg d_A$

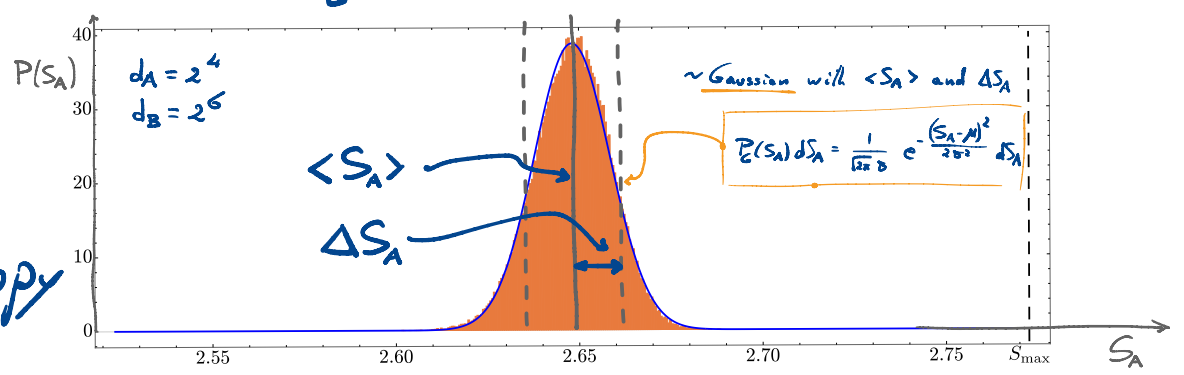
■ Variance [Bianchi-Donà, PRD'19]

$$(\Delta S_A)^2 = \langle S_A^2 \rangle - \langle S_A \rangle^2$$

$$(\Delta S_A)^2 = \frac{d_A + d_B}{d_A d_B + 1} \Psi'(d_B + 1) - \Psi'(d_A d_B + 1) - \frac{(d_A - 1)(d_A + 2d_B - 1)}{4 d_B^2 (d_A d_B + 1)}$$

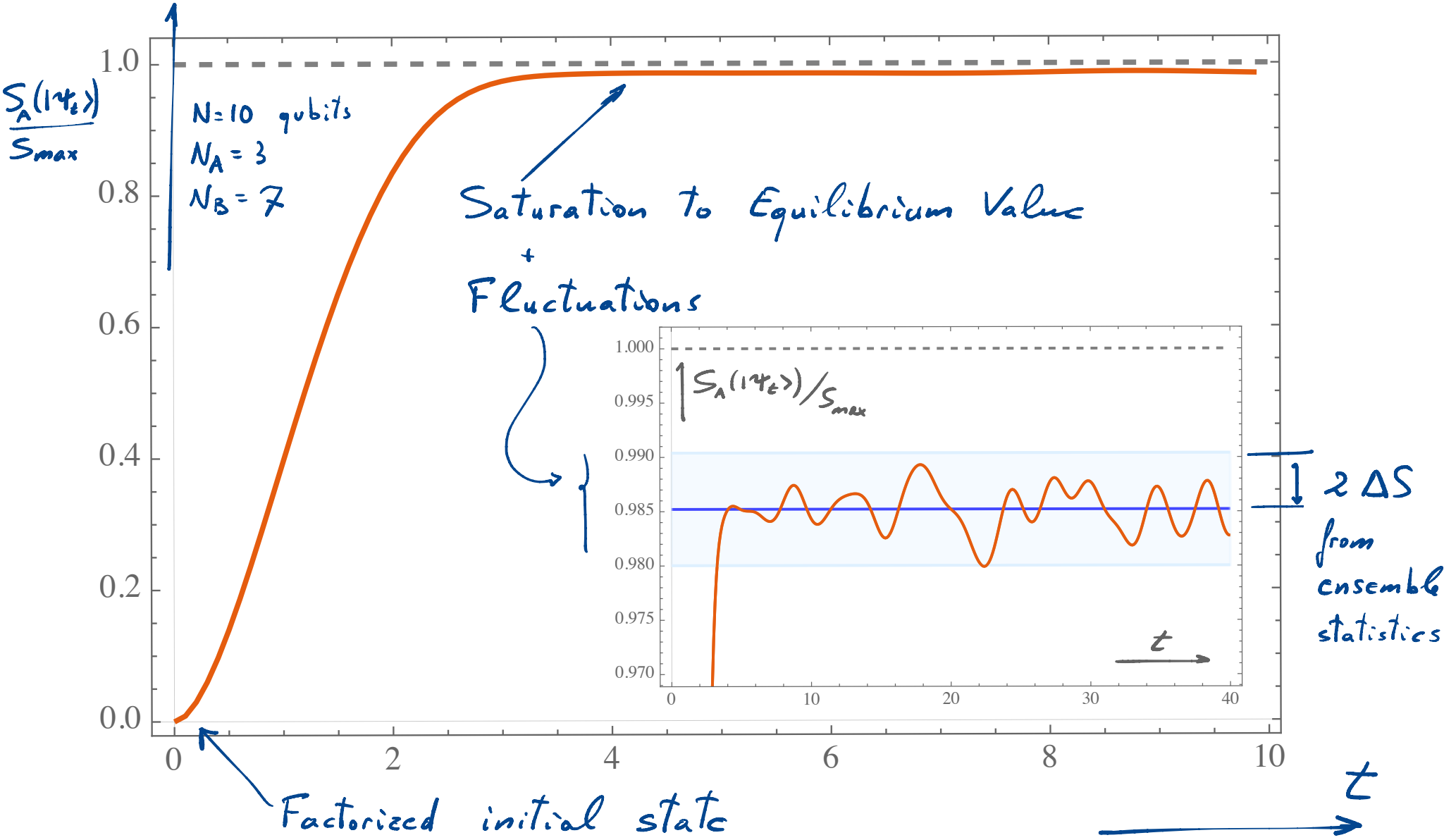
• Asymptotics: $(\Delta S_A)^2 \sim \frac{1}{2 d_B^2}$ narrow distribution
 $d_B \gg d_A$

⇒ Typical value of the entanglement entropy



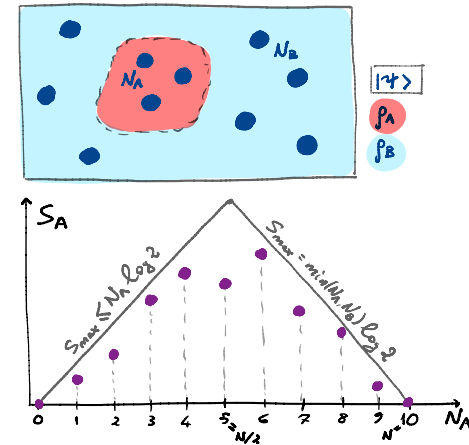
Entanglement Entropy and Thermalization

- Random factorized initial state $|\psi_0\rangle = |\phi_A\rangle|\chi_B\rangle$ (quantum quench)
- Evolution with chaotic Hamiltonian, $|\psi_t\rangle = e^{-iHt}|\psi_0\rangle$



Typical entanglement with constraints

Page '93: qubit model of unitary BH evaporation
 * note: no Hamiltonian



* Assumptions:

- Finite dimension $\dim \mathcal{H}$

| | | |
|---|--------|---|
| } | qubits | ✓ |
| | QFT | x |
| | QG | ? |

- Factorization $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$

| | | |
|---|--------|------------------------|
| } | qubits | ✓ |
| | QFT | x (Reeh-Schlieder Thm) |
| | QG | x |

Constraints: $G^i |\psi\rangle = 0$

$$\mathcal{H}^{(j)} = \bigoplus_{\ell} \left(\mathcal{H}_A^{(\ell)} \otimes \mathcal{H}_{\bar{A}}^{(j,\ell)} \right)$$

finite dim, non-factorized

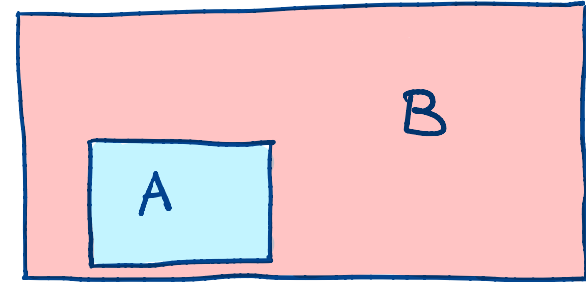
eg.:

- qubits → fixed N excitations
- cold atoms in optical lattice → N atoms
- QFT → fixed energy
- QG → Diff & H constraints
- BH mass → fixed energy

□ Typical entanglement with constraints : abelian example

• Hilbert Space $\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$

• Hamiltonian $H = H_A + H_B$ with $[H_A, H_B] = 0$



• Basis $|\epsilon_A, \alpha\rangle |\epsilon_B, \beta\rangle$, energy eigenspace $E = \epsilon_A + \epsilon_B$

• Direct sum decomposition and energy conservation

$$\mathcal{H} = \bigoplus_E \mathcal{H}^{(E)}$$

$$\mathcal{H}^{(E)} = \bigoplus_{\epsilon_A} \left(\mathcal{H}_A^{(\epsilon_A)} \otimes \mathcal{H}_B^{(E - \epsilon_A)} \right)$$

• Energy eigenstates and block-diagonal density matrix

$$H |\psi_E\rangle = E |\psi_E\rangle \quad \Rightarrow \quad [\rho_A, H_A] = 0$$

with $\rho_A = \text{Tr}_B |\psi_E\rangle \langle \psi_E|$

$$\Rightarrow \quad \rho_A = \begin{pmatrix} \square & & & \\ & \square & & \\ & & \square & \\ & & & \dots \end{pmatrix} = \bigoplus_{\epsilon_A} p(\epsilon_A) \rho(\epsilon_A)$$

• Entanglement entropy

$$S_A(|\psi_E\rangle) \equiv -\text{Tr}(\rho_A \log \rho_A) = \sum_{\epsilon_A} p(\epsilon_A) (-\text{Tr} \rho(\epsilon_A) \log \rho(\epsilon_A)) - \sum_{\epsilon_A} p(\epsilon_A) \log p(\epsilon_A)$$

Exact formulas for:

$$\mathcal{H}^{(E)} = \bigoplus_r \left(\mathcal{H}_A^{(r)} \otimes \mathcal{H}_{\bar{A}}^{(r)} \right)$$

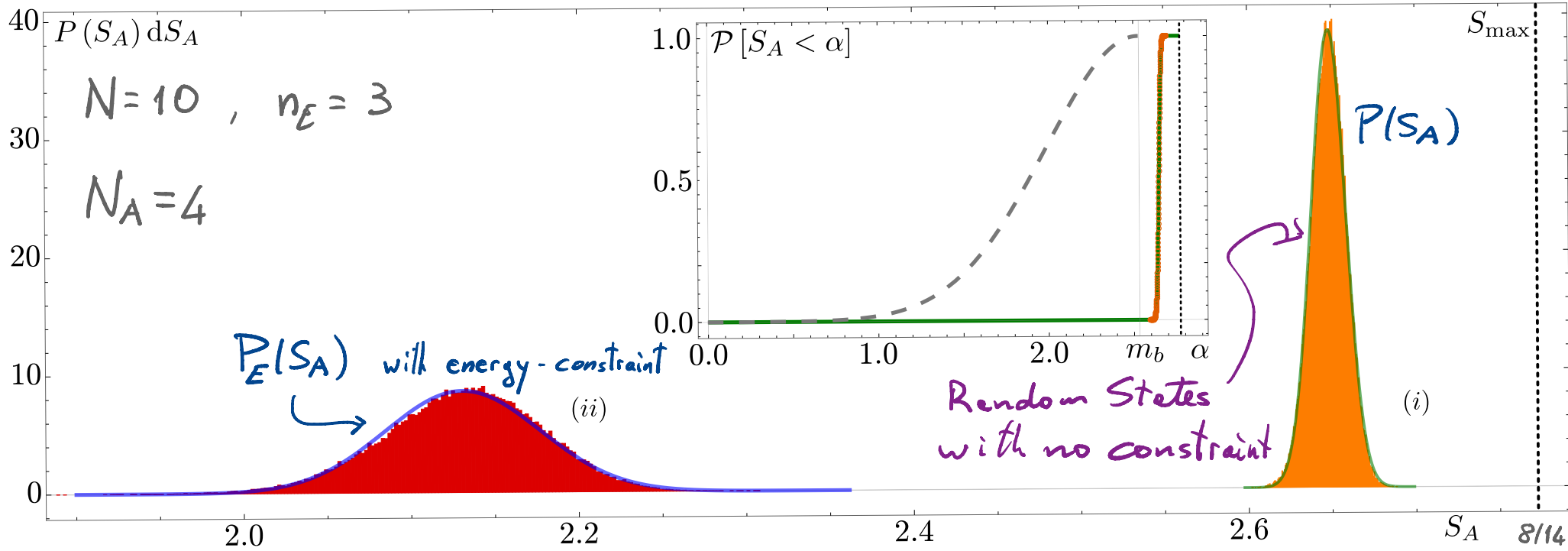
$$d_r = \dim \mathcal{H}_A^{(r)} \quad , \quad b_r = \dim \mathcal{H}_{\bar{A}}^{(r)} \quad , \quad D = \sum_r d_r b_r$$

• Average entanglement entropy

$$\langle S_A \rangle_E = \sum_r \frac{d_r b_r}{D} \left(\Psi(D+1) - \Psi(\max(d_r, b_r) + 1) - \min\left(\frac{d_r-1}{2b_r}, \frac{b_r-1}{2d_r}\right) \right)$$

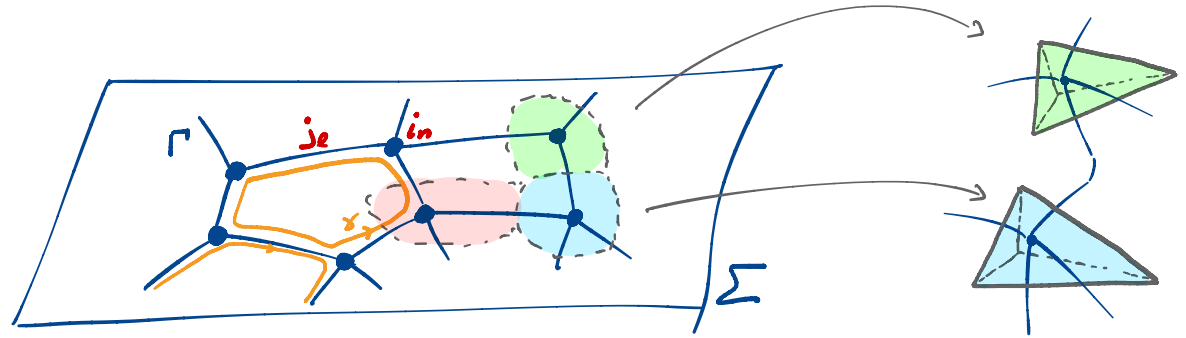
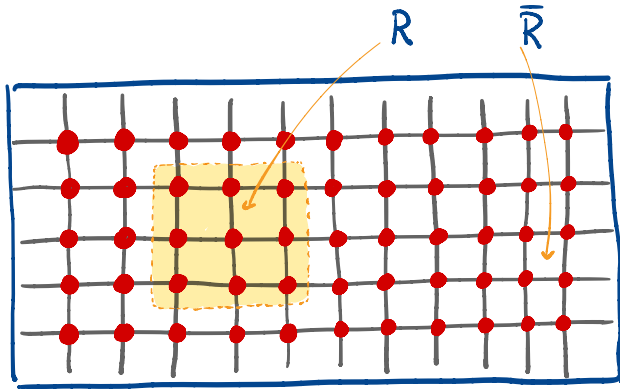
• Variance

$$(\Delta S)_E = \dots$$



② Hierarchy of states: Volume law, Area law, Zero law

• Many-body system $\mathcal{H} = \bigotimes_{n=1}^N \mathcal{H}_n = \mathcal{H}_R \otimes \mathcal{H}_{\bar{R}}$ vs LQG



Ⓐ Volume-Law States

- Random States
- High Temperature

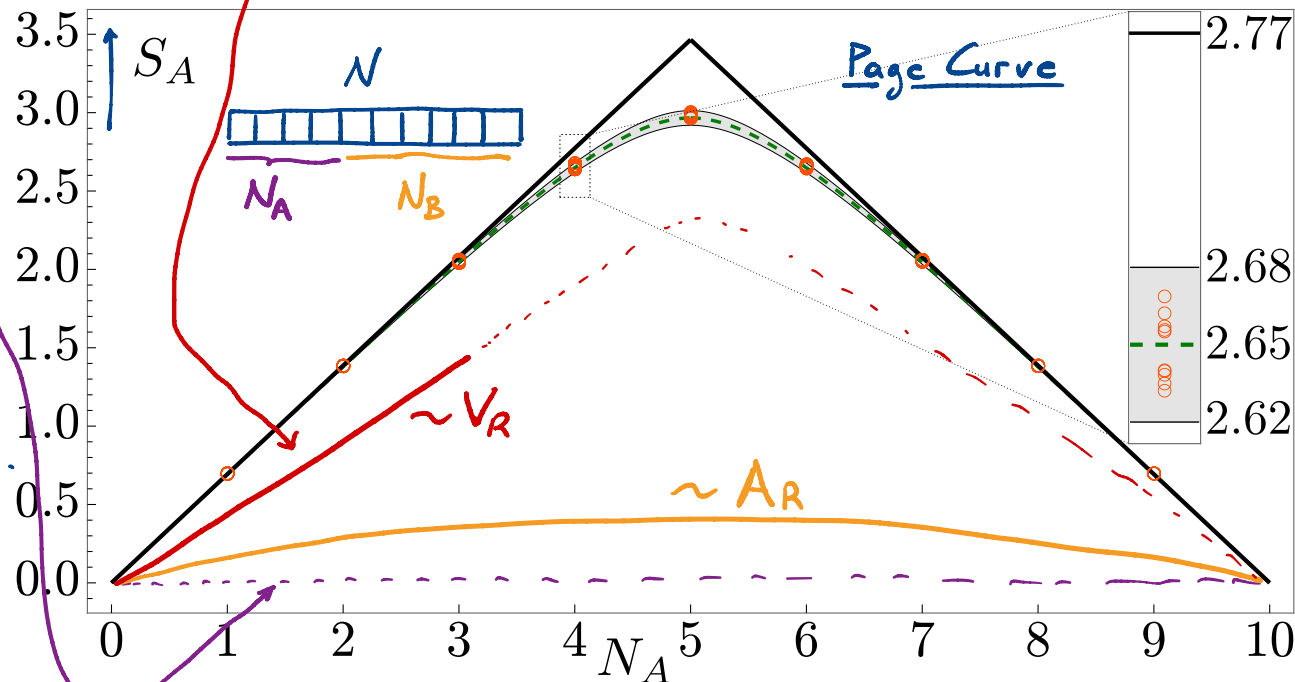
$S_R(|\psi\rangle) = \delta V_R + \dots$

Ⓑ Zero-Law States $S_R(|\psi\rangle) = c_0 + \dots$

- QG basis states
- High Energy (no T)

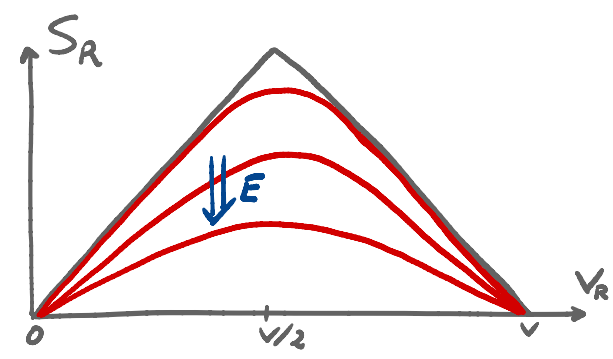
Ⓒ Area-Law States $S_R(|\psi\rangle) = \alpha A_R + \dots$

- Ground State of Local H
- Long-Range Correlations



□ Area Law & Semiclassicality

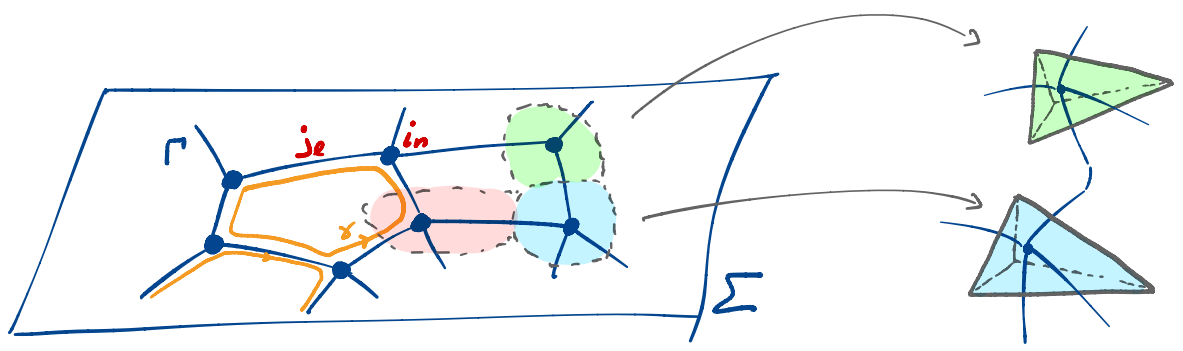
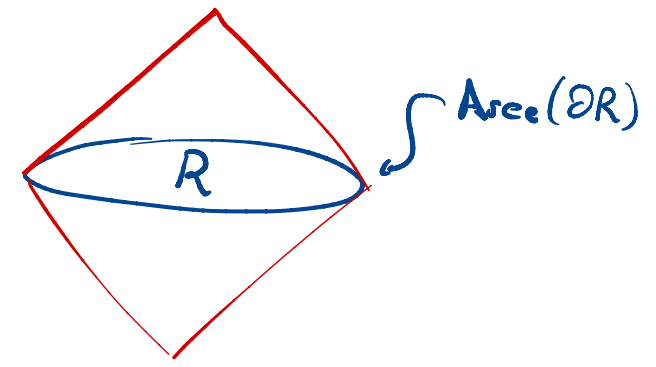
- In CM & QFT, as we lower the energy, we transition from volume-law to area-law
- In CM, zero law states are high-energy (not Fock in QFT)
- In QG, we don't have an immediate notion of energy or energy-density



⇒ Reverse Perspective: Entanglement as a Probe

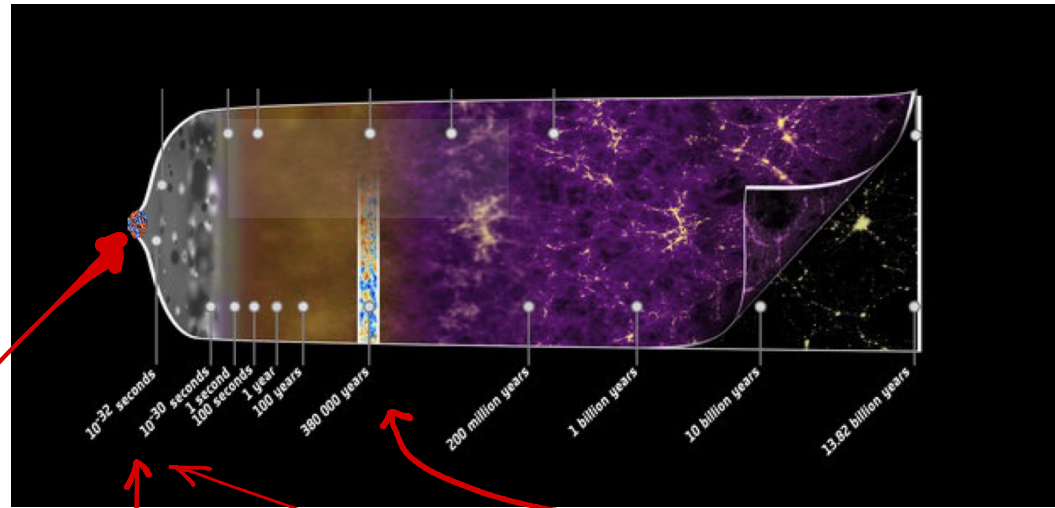
Architecture Conjecture (Bianchi-Myers [1212.5183] (QG))

Semiclassical $|\psi\rangle$ in QG belong to the area-law corner of $\mathcal{H}_{\text{phys}}$

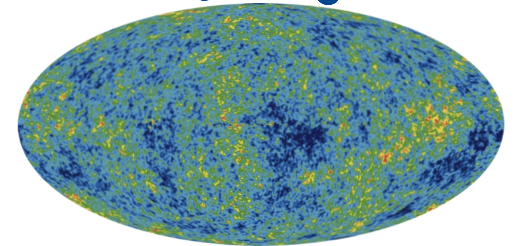
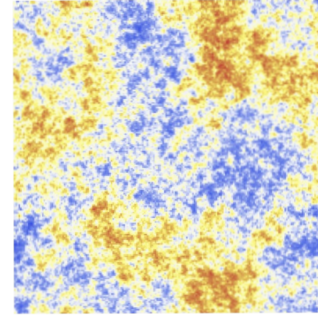
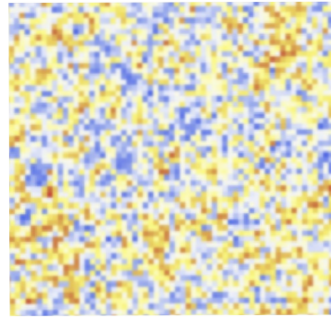
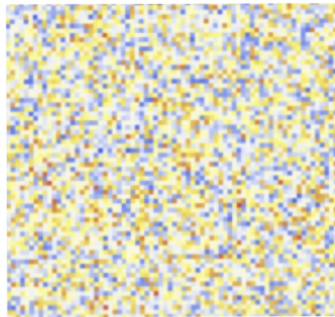
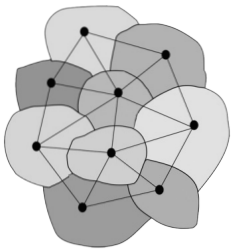
$$S_R(|\psi\rangle) = 2\pi \frac{\langle \text{Area}(\partial R) \rangle}{\ell_p^2} + \dots$$


Volume Law and Zero Law States are genuine quantum geometries far from classical spacetime + quantum perturbation

3 Zero law and causal origin of entanglement

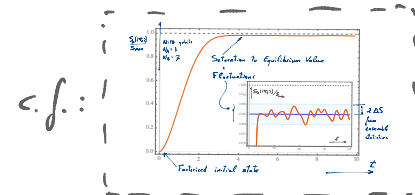


Planck scale → Pre-Inflationary Phase → Inflation → Hot Big Bang & CMB

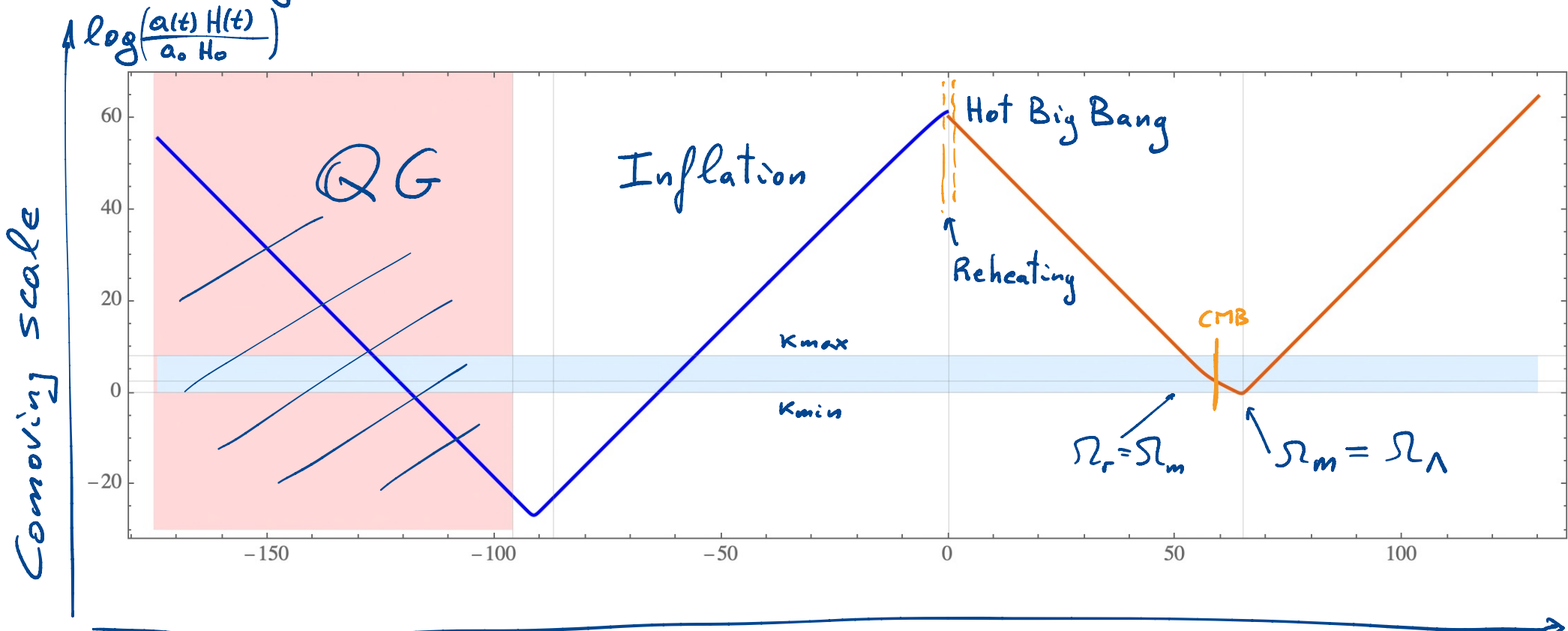


$| \Psi_0 \rangle$

Zero-law state, cosmological quench
and quantum BKL conjecture



□ Entanglement Dynamics in the Primordial Universe



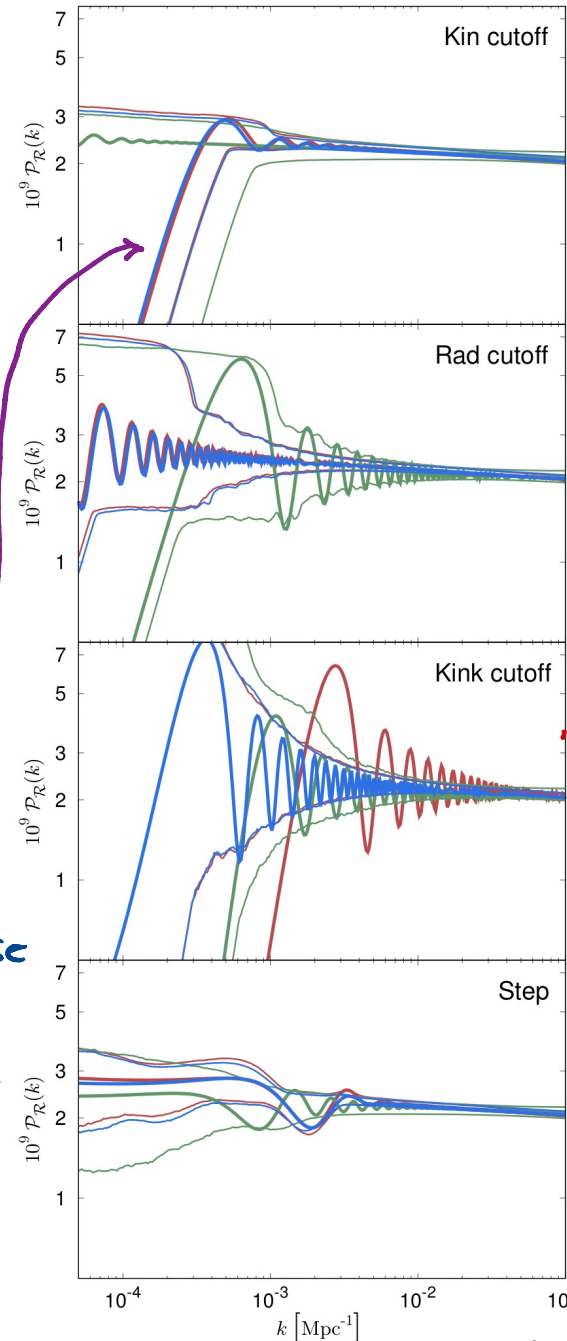
$n_{\text{foldings}} = N = \log \frac{a(t)}{a_{\text{HBB}}}$

- Inflation → Gaussian Squeezing
- Before Inflation → QG-to-QFT Transition

Power Spectrum & Initial State of Perturbations

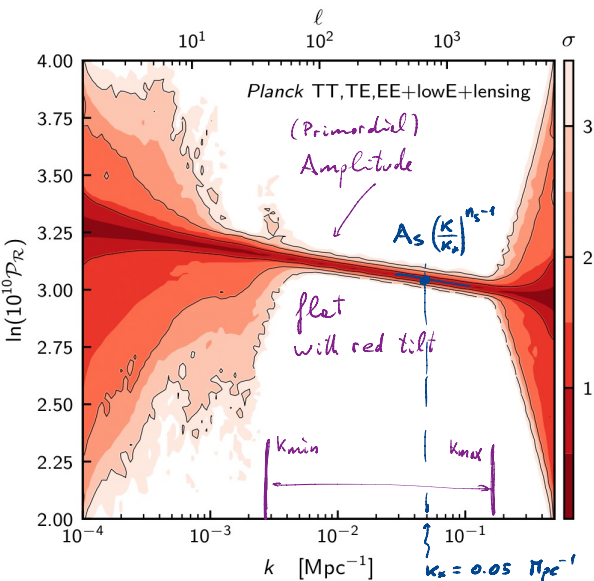
Planck 2018 results. X. Constraints on inflation [1807.06211]

- Initial state $|4_0\rangle$: Bunch-Davies vacuum vs excited
- Pre-inflationary phase can prepare excited state
- If inflation doesn't last too long ($N_{\text{infl}} \sim 60$)
 \Rightarrow relic of the initial state
 at large scales (at low k)
- Zero-law state \Rightarrow power suppression



* Note:

- degenerate with kinetic phase
- PLANCK 2018 low- l anomaly
- Statistical significance?
- Other observables?



□ Lec. 1 LQG, ref: Ashtekar-Bianchi, "A short review..." Rep. Prog. Phys 2021 [2104.04394]

- Area is more fundamental than length
- Loops are more fundamental than curvature
- QG as TQFT with 2d defects

□ Lec. 2 LQG & QI, ref: Bianchi-Livine, in "Handbook of QG" Springer 2023 [2302.05922]

- Hierarchy of entanglement in CMT and QFT
- Area law in LQG
- Primordial entanglement