



Kinetic theory of massive spin-1/2 particles from the Wigner-function formalism

N. Weickgenannt, X.-I. Sheng, E. Speranza, Q. Wang, and D.H. Rischke, arXiv:1902.06513

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Why magneto-hydrodynamics (MHD)?



 Early stage of non-central heavy-ion collisions: large orbital angular momenta and strong electromagnetic fields.

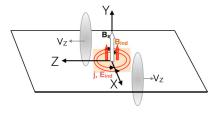


Figure from V. Roy, S. Pu, L. Rezzolla, and D. H. Rischke, PRC96 (2017) 054909

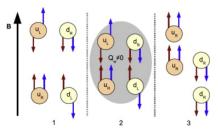
- Strong electromagnetic fields in early universe and compact stars.
- For massive spin-0 particles, second-order dissipative MHD has already been studied.
 - G.S. Denicol, X.-G. Huang, E. Molnar, G.M. Monteiro, H. Niemi, J. Noronha, D.H. Rischke, and Q. Wang, PRD 98 (2018) 076009
 - G.S. Denicol, E. Molnar, H. Niemi, and D.H. Rischke, PRD 99 (2019) 056017
- But all elementary matter particles are fermions...



Spin effects in heavy-ion collisions



- Chiral vortical effect (CVE): charge currents induced by vorticity.
- Chiral magnetic effect (CME): charge currents induced by magnetic fields.



Dmitri E. Kharzeev, Larry D. McLerran, and Harmen J. Warringa, NPA 803 (2008)

Has been studied in massless case.

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J-Y. Chen, D. T. Son, and M. Stephanov, PRL 115 (2015), 021601;
Y. Hidaka, S. Pu, and D-L. Yang, PRD95 (2017) 091901;
A. Huang, S. Shi, Y. Jiang, J. Liao, and P. Zhuang, PRD 98 (2018) 036010
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Similar effects for massive particles?





■ What we want: kinetic theory and fluid dynamics for massive spin-1/2 particles in inhomogeneous electromagnetic fields.

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    J.-H. Gao, and Z.-T. Liang, arXiv:1902.06510 (2019)
    K. Hattori, Y. Hidaka, and D.-L. Yang, arXiv:1903.01653 (2019)
    Z. Wang, X. Guo, S. Shi, and P. Zhuang, arXiv:1903.03461 (2019)
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- Starting point: quantum field theory, Dirac equation.
- Strategy: use Wigner functions to derive kinetic theory.
- Goal: determine fluid-dynamical equations of motion from resulting Boltzmann equation.

Wigner functions

- Quantum analogue of classical distribution function.
- Contains information about quantum state of system.
- Off-equilibrium: two-point function depends not only on relative coordinate y, but also on central coordinate x.
- Wigner transformation of two-point function:
 H.-Th. Elze, M. Gyulassy, and D. Vasak, Ann. Phys. 173 (1987) 462

$$W(x,p) = \int \frac{d^4 y}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar}p \cdot y} \langle : \bar{\Psi}(x+\frac{y}{2})\Psi(x-\frac{y}{2}) : \rangle,$$

- Quantum analogue of classical distribution function.
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 H.-Th. Elze, M. Gyulassy, and D. Vasak, Ann. Phys. 173 (1987) 462

$$W(x,p) = \int \frac{d^4 y}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar}p \cdot y} \left\langle : \bar{\Psi}(x + \frac{y}{2})U(x + \frac{y}{2}, x)U(x, x - \frac{y}{2})\Psi(x - \frac{y}{2}) : \right\rangle,$$

with gauge link

$$U(b,a) \equiv P \exp \left(-rac{i}{\hbar} \int_a^b dz^\mu A_\mu(z)
ight)$$

to ensure gauge invariance.

Calculating the Wigner function



- In general: result of calculation of Wigner function directly from definition is not on-shell.
- Momentum variable of directly calculated Wigner function is physical (kinetic) momentum only at zeroth order in \hbar/g radients.
- Dirac equation implies transport equation for Wigner function.
- Idea: Find general solutions for this transport equation by expanding in powers of \hbar .
- lacksquare Decompose W in transport equation into generators of Clifford algebra:

$$W = \frac{1}{4} \left(\mathcal{F} + i \gamma^5 \mathcal{P} + \gamma^{\mu} \mathcal{V}_{\mu} + \gamma^5 \gamma^{\mu} \mathcal{A}_{\mu} + \frac{1}{2} \sigma^{\mu\nu} \mathcal{S}_{\mu\nu} \right).$$

- Equations for \mathcal{F} (scalar, "distribution function") and $\mathcal{S}_{\mu\nu}$ (tensor, "dipole moment") decouple from rest.
- Determine V_{μ} ("vector current"), A_{μ} ("polarization"), \mathcal{P} from $S_{\mu\nu}$, \mathcal{F} .
- Results will hold up to order $\mathcal{O}(\hbar)$.
- Notation: $W = W^{(0)} + \hbar W^{(1)} + \mathcal{O}(\hbar^2)$.



General results



$$\begin{split} \mathcal{F} &= m \left[V \, \delta(\rho^2 - m^2) - \frac{\hbar}{2} F^{\mu\nu} \Sigma^{(0)}_{\mu\nu} A^{(0)} \, \delta'(\rho^2 - m^2) \right] + \mathcal{O}(\hbar^2) \,, \\ \mathcal{S}_{\mu\nu} &= m \left[\bar{\Sigma}_{\mu\nu} \delta(\rho^2 - m^2) - \hbar F_{\mu\nu} V^{(0)} \delta'(\rho^2 - m^2) \right] + \mathcal{O}(\hbar^2) \,, \\ \mathcal{P} &= \frac{\hbar}{4m} \epsilon^{\mu\nu\alpha\beta} \nabla_{\mu} \left[p_{\nu} \Sigma^{(0)}_{\alpha\beta} A^{(0)} \, \delta(\rho^2 - m^2) \right] + \mathcal{O}(\hbar^2) \,, \\ V_{\mu} &= \delta(\rho^2 - m^2) \left[p_{\mu} V + \frac{\hbar}{2} \nabla^{\nu} \Sigma^{(0)}_{\mu\nu} A^{(0)} \right] \\ &- \hbar \left[\frac{1}{2} p_{\mu} F^{\alpha\beta} \Sigma^{(0)}_{\alpha\beta} + \Sigma^{(0)}_{\mu\nu} F^{\nu\alpha} p_{\alpha} \right] A^{(0)} \, \delta'(\rho^2 - m^2) + \mathcal{O}(\hbar^2) \,, \\ \mathcal{A}_{\mu} &= -\frac{1}{2} \epsilon_{\mu\nu\alpha\beta} \rho^{\nu} \bar{\Sigma}^{\alpha\beta} \, \delta(\rho^2 - m^2) + \hbar \tilde{F}_{\mu\nu} \rho^{\nu} V^{(0)} \, \delta'(\rho^2 - m^2) + \mathcal{O}(\hbar^2) \,, \end{split}$$

with

$$\begin{array}{lcl} \bar{\Sigma}^{(0)\mu\nu} & = & \Sigma^{(0)\mu\nu} A^{(0)}, \\ \rho_{\nu} \bar{\Sigma}^{\mu\nu} & = & \hbar \nabla^{\mu} V^{(0)} \\ \nabla^{\mu} & \equiv & \partial^{\mu}_{\nu} - F^{\mu\nu} \partial_{\nu\nu} \end{array}$$

Boltzmann equation for massive spin-1/2 particles



Unknown $V=f_++f_-$ and $A=f_+-f_-$ are determined by generalized Boltzmann equation

$$\sum_s \delta \left(p^2 - m^2 - \frac{s}{2} \hbar F^{\alpha\beta} \Sigma^{(0)}_{\alpha\beta} \right) \left\{ p^\mu \partial_{x\mu} f_s + \partial_{p\mu} \left[F^{\mu\nu} p_\nu + \frac{\hbar}{4} s \Sigma^{(0)\nu\rho} (\partial^\mu F_{\nu\rho}) \right] f_s \right\} = 0$$

- f_s distribution functions for spin-up (s = +) and spin-down (s = -) particles.
- Modified on-shell condition!
- Force on particle: first Mathisson-Papapetrou-Dixon (MPD) equation \rightarrow Particle with classical dipole moment $\Sigma^{(0)\mu\nu}$ in electromagnetic field:
 - W. Israel, General Relativity and Gravitation 9 (1978) 451

$$mrac{d}{d au}p^{\mu}=F^{\mu
u}p_{
u}+rac{\hbar}{4}s\Sigma^{(0)
u
ho}(\partial^{\mu}F_{
u
ho})\,.$$

au: worldline parameter, $rac{d}{d au}=\dot{\chi}^{\mu}rac{\partial}{\partial x^{\mu}}+\dot{p}^{\mu}rac{\partial}{\partial p^{\mu}}.$

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Dipole moment and classical limit

- $ar{oldsymbol{\Sigma}}_{\mu
 u}$ determined by kinetic equation for dipole moment.
- To zeroth order:

$$m\frac{d}{d\tau}\Sigma^{(0)\mu\nu} = \Sigma^{(0)\lambda\nu}F^{\mu}_{\ \lambda} - \Sigma^{(0)\lambda\mu}F^{\nu}_{\ \lambda},$$

- Recover second MPD equation for dipole-moment tensor $\Sigma_{\mu\nu}^{(0)}$!

 W. Israel, General Relativity and Gravitation 9 (1978) 451
- Equivalent to Bargmann-Michel-Telegdi (BMT) equation
 V. Bargmann, L. Michel, and V.L. Telegdi, PRL 2 (1959) 435

$$m\frac{d}{d\tau}n^{(0)\mu}=F^{\mu\nu}n^{(0)}_{\nu},$$

with classical spin vector

$$n^{(0)\mu} = -\frac{1}{2m} \epsilon^{\mu\nu\alpha\beta} p_{\nu} \Sigma^{(0)}_{\alpha\beta}$$



Non-relativistic dipole-moment tensor connected to spin three-vector n^k :

$$\Sigma^{ij} = \epsilon^{ijk} n^k.$$

- For massive particles: define spin in rest frame.
 - U. Heinz, PLB 144 (1984) 228

$$\Sigma^{\mu
u} = -rac{1}{m}\epsilon^{\mu
ulphaeta}p_{lpha}{}^{}{}_{eta}.$$

- For massless particles: define spin in arbitrary frame with four-velocity u^{μ} . Spin vector n^{μ} is always parallel to momentum.
 - J-Y. Chen, D.T. Son, and M. Stephanov, PRL 115 (2015) 021601

$$\Sigma_{u}^{\mu\nu} = -rac{1}{oldsymbol{p}\cdotoldsymbol{u}}\epsilon^{\mu
ulphaeta}oldsymbol{u}_lphaoldsymbol{p}_eta.$$

- Massless limit: replace massive by massless dipole-moment tensor $\Sigma^{\mu\nu} \to \Sigma^{\mu\nu}_n$.
- Result agrees with previously known massless solution!
 - Y. Hidaka, S. Pu, and D-L. Yang, PRD 95 (2017) 091901
 - A. Huang, S. Shi, Y. Jiang, J. Liao, and P. Zhuang, PRD 98 (2018) 036010
 - J.-H. Gao, Z.-T. Liang, Q. Wang, and X.-N. Wang, PRD 98 (2018) 036019

Vector current in global equilibrium



- In global equilibrium: Analytic solution for Boltzmann equation.
- Vector current is explicitly calculated as:

$$\begin{split} \mathcal{V}^{\mu} &= \frac{2}{(2\pi\hbar)^{3}} \sum_{s} \left[\delta(\rho^{2} - m^{2}) \left(\rho^{\mu} - m\hbar \frac{s}{2} \tilde{\omega}^{\mu\nu} n_{\nu}^{(0)} \partial_{\beta \cdot \pi} \right) \right. \\ &+ \left. \hbar s \tilde{F}^{\mu\nu} n_{\nu}^{(0)} \delta'(\rho^{2} - m^{2}) + \hbar \frac{s}{2m} \delta(\rho^{2} - m^{2}) \epsilon^{\nu\mu\alpha\beta} \rho_{\alpha} \nabla_{\nu} n_{\beta}^{(0)} \right] f_{s}^{(0)} \,, \end{split}$$

with zeroth-order equilibrium distribution function

$$f_s^{(0)} = [\exp(\beta \cdot \pi - \beta \mu_s) + 1]^{-1},$$

where π^{μ} canonical momentum, β^{μ} thermal fluid velocity, β inverse temperature, μ_s chemical potential.

- Analogue of chiral vortical effect (CVE) for massive particles.
 - D. T. Son and P. Surowka, PRL 103 (2009) 0906.5044
- Analogue of chiral magnetic effect (CME).
 - D. E. Kharzeev, L. D. McLerran, and H. J. Warringa, NPA 803 (2008) 0711.0950
- Thermal vorticity tensor: $\omega_{\mu\nu} \equiv \frac{1}{2} (\partial_{\mu}\beta_{\nu} \partial_{\nu}\beta_{\mu}).$
- Dual thermal vorticity tensor: $\tilde{\omega}_{\mu\nu} \equiv \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} \omega^{\alpha\beta}$.



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Axial-vector current in global equilibrium

Obtain expression for axial-vector current:

$$\mathcal{A}^{\mu} = \frac{2}{(2\pi\hbar)^3} \sum_{s} \left[\delta(\rho^2 - m^2) \left(s \, m \, n^{(0)\mu} - \frac{\hbar}{2} \tilde{\omega}^{\mu\nu} \rho_{\nu} \partial_{\beta \cdot \pi} \right) \right.$$
$$\left. + \, \hbar \tilde{F}^{\mu\nu} \rho_{\nu} \delta'(\rho^2 - m^2) \right] f_s^{(0)} - \frac{\hbar}{2} \epsilon^{\mu\nu\alpha\beta} \rho_{\nu} \Xi_{\alpha\beta} \, \delta(\rho^2 - m^2) \, .$$

- Classical spin precession.
- Analogue of axial chiral vortical effect (ACVE).
- Analogue of chiral separation effect (CSE).
- D. E. Kharzeev, J. Liao, S. A. Voloshin, and G. Wang, Prog. Part. NP 88 (2016), 1511.04050

Fluid-dynamical equations with spin I



■ Particle current:

$$J^{\mu}=\int d^4p\,\mathcal{V}^{\mu}.$$

Not parallel to the fluid velocity!

$$\partial_{\mu}J^{\mu}=0.$$

Conserved!

Canonical energy-momentum tensor (matter part):

$$T^{\mu\nu}_{mat} = \int d^4p \, p^{
u} \mathcal{V}^{\mu}.$$

Not symmetric!

$$\partial_{\mu} T_{mat}^{\mu\nu} = F^{\nu\mu} J_{\mu}.$$

Conserved in combination with electromagnetic and interaction part.

Fluid-dynamical equations with spin II



■ Canonical spin tensor (matter part):

$$\mathcal{S}_{\it mat}^{\lambda,\mu
u} = -rac{1}{2}\epsilon^{\mu
u\lambda
ho}\int d^4p\,\mathcal{A}_
ho$$

$$\hbar\partial_{\lambda}S_{mat}^{\lambda,\mu
u}=T_{mat}^{
u\mu}-T_{mat}^{\mu
u}$$

Not conserved!

Spin angular momentum and orbital angular momentum are converted into one another.

- ightarrow Consideration of spin leads to additional fluid-dynamical equation of motion.
- W. Florkowski, B. Friman, A. Jaiswal, and E. Speranza, PRC 97 (2018) 041901;
- W. Florkowski, B. Friman, A. Jaiswal, R. Ryblewski, and E. Speranza, PRD 97 (2017) 116017;
- W. Florkowski, F. Becattini, and E. Speranza, APB 49 (2018) 1409:
- F. Becattini, W. Florkowski, and E. Speranza, PLB 789 (2019) 419-425

Conclusions



- Derived transport equation for distribution function and polarization for massive spin-1/2 particles in inhomogeneous electromagnetic fields.
- Recovered classical equations of motion.
- Solution agrees with previously known massless solution in massless limit.
- Derived explicit expressions for currents in global equilibrium.
- Found analogues of CVE, CME, ACVE, and CSE for massive particles.
- Derived fluid-dynamical equations of motion.

Outlook



- Solve generalized Boltzmann equation.
- Include collisions.
 - → Boltzmann equation without assuming equilibrium.
- Derive equations of motion for dissipative quantities.
 - \rightarrow Method of moments.

Back-up

Conventions and Definitions



- Natural units, $c = k_B = 1$, but keep \hbar explicitly.
- To simplify notation: only write positive-energy parts of solutions.
- Polarization direction n^{μ} : space-like unit vector parallel to axial-vector current.
- Spin quantization direction: unit vector, purely spatial in particle rest frame.

"spin up", s=+: projection of spin onto quantization direction positive. "spin down", s=-: projection of spin onto quantization direction negative. Here: chosen to be identical to polarization direction.

$$\bar{\textit{u}}_{\textit{s}} \gamma^{\mu} \gamma^{\textit{5}} \textit{u}_{\textit{s}} = 2 \textit{ms} \; \textit{n}^{\mu}$$

Spin tensor vs. dipole-moment tensor

■ Dipole-moment tensor:

$$s\Sigma^{\mu\nu} = \frac{1}{2m}\bar{u}_s \frac{i}{2} [\gamma^{\mu}, \gamma^{\nu}] u_s$$
$$= -s \frac{1}{m} \epsilon^{\mu\nu\alpha\beta} p_{\alpha} n_{\beta}$$

called "spin tensor" in

U. Heinz, PLB 144 (1984) 228,

J-Y. Chen, D. T. Son, and M. Stephanov, PRL 115 (2015), 021601

Y. Hidaka, S. Pu, D-L. Yang, PRD 95 (2017) 091901

S.R. De Groot, Relativistic Kinetic Theory. Principles and Applications (1980)

Spin tensor:

rank-3 tensor $\mathcal{S}^{\lambda,\mu
u}$ such that total angular momentum

$$J^{\lambda,\mu\nu} = x^{\mu} T^{\lambda\nu} - x^{\mu} T^{\lambda\mu} + \hbar S^{\lambda,\mu\nu}$$

Diagonal spin basis

- Distribution function f_{rs} is Hermitian matrix in spin space.
- Can be diagonalized by Unitary transformation:

$$f_{rs} = D_{rr'} \tilde{f}_{s'} \delta_{r's'} D_{s's}^{\dagger}.$$

Redefine spinors

$$\tilde{u}_s \equiv \sum_{s'} u_{s'} D_{s's}.$$

Define

$$\mathit{sn}^{\mu} \equiv \overline{\tilde{\mathit{u}}}_{\mathit{s}} \gamma^{\mu} \gamma^{\mathit{5}} \mathit{u}_{\mathit{s}}.$$

Only diagonal part contributes!

Transport equation

From Dirac equation: transport equation for Wigner function:

H.-Th. Elze, M. Gyulassy, and D. Vasak, Ann. Phys. 173 (1987) 462

$$(\gamma_{\mu}K^{\mu}-m)W(X,p)=0,$$

with

$$K^{\mu} \equiv \Pi^{\mu} + \frac{1}{2}i\hbar\nabla^{\mu},$$

$$\nabla^{\mu} \equiv \partial_{x}^{\mu} - j_{0}(\Delta)F^{\mu\nu}\partial_{\rho\nu},$$

$$\Pi^{\mu} \equiv \rho^{\mu} - \hbar\frac{1}{2}j_{1}(\Delta)F^{\mu\nu}\partial_{\rho\nu},$$

 $\Delta = \frac{1}{2}\hbar\partial_{\rho}\cdot\partial_{\kappa}$ with ∂_{κ} only acting on $F^{\mu\nu}$ and $j_0(r)=\sin(r)/r$, $j_1(r) = [\sin(r) - r\cos(r)]/r^2$ spherical Bessel functions.

- Exact quantum kinetic equation for Wigner function for massive spin 1/2-particles and inhomogeneous fields!
- Only assumption: vanishing collision kernel, external classical gauge fields.

Massless limit: details



- $lacksquare\ p\cdot u$ related to rest-frame energy $\sqrt{p^2} o\delta$ -function!
- Massless limit: replace massive by massless dipole-moment tensor $\Sigma^{\mu\nu} o \Sigma^{\mu\nu}_u$.
- Attention: $\delta(p^2 m^2)/m \rightarrow \delta(p^2)/(p \cdot u)$.
- ullet Find general solution for constraint on $ar{\Sigma}^{\mu
 u}$.
- lacksquare Define right- and left-handed currents $J^\pm_\mu\equivrac{1}{2}({\cal V}^{m=0}_\mu\pm{\cal A}^{m=0}_\mu)$
- Result

$$J_{\mu}^{\pm} = \left[p_{\mu} \delta(\rho^2) \pm \frac{1}{2} \hbar \epsilon_{\mu\nu\alpha\beta} \rho^{\nu} F^{\alpha\beta} \delta'(\rho^2) \pm \frac{1}{2} \hbar \epsilon_{\mu\nu\alpha\beta} \frac{\rho^{\alpha} u^{\beta}}{\rho \cdot u} \delta(\rho^2) \nabla^{\nu} \right] f_{\pm}.$$

agrees with previously known massless solution!

Y. Hidaka, S. Pu, D-L. Yang, PRD 95 (2017) 091901; A. Huang, S. Shi, Y. Jiang, J. Liao, and P. Zhuang, PRD 98 (2018) 036010; J.-H. Gao, Z.-T. Liang, Q. Wang, and X.-N. Wang, PRD 98 (2018) 036019

Distribution functions in global equilibrium



Equilibrium distribution function:

$$f_s^{eq} = (e^{g_s} + 1)^{-1},$$

with g linear combination of conserved quantities charge, momentum, and angular momentum:

$$g_s = \beta \pi \cdot U - \beta \mu_s + \frac{\hbar}{4} s \Sigma^{\mu\nu} \partial_{\mu} (\beta U_{\nu}).$$

Here, $\pi_{\mu} \equiv p_{\mu} + A_{\mu}$ is canonical momentum, U_{μ} is fluid velocity, $eta \equiv rac{1}{ au}$ is inverse temperature, and μ_s is chemical potential.

To zeroth order

$$f_s^{(0)} = (e^{g_{s0}} + 1)^{-1},$$

with

$$g_{s0} = \beta(\pi \cdot U - \mu_s).$$



Equilibrium conditions



■ "Homogeneous" part of the Boltzmann equation fulfilled if:

$$\begin{array}{rcl} \mu_{\text{s}} & = & \text{const}, \\ \partial_{\nu}\beta_{\mu} + \partial_{\mu}\beta_{\nu} & = & 0, \end{array}$$

 "Inhomogeneous" part of Boltzmann equation: additional conditions to make global equilibrium possible, e.g.

$$\mu_{s=+} - \mu_{s=-} = 0$$
 or $\partial_{x\alpha} F_{\mu\nu} = 0$.